

# Bose-Einstein Condensation (BEC)

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occurs when  $z = 1 \Rightarrow \mu = 0$

$$\langle n \rangle \lambda_T^3 = g_{3/2}(1) \approx 2.612$$

determines phase diagram: \* for a given temperature  $T$  density at which BEC occurs

$$\text{is } \langle n \rangle_c \lambda_T^3 = g_{3/2}(1)$$

\* for a given

density  $\langle n \rangle$ , temperature at which BEC occurs is  $\langle n \rangle \lambda_{T_c}^3 = g_{3/2}(1)$

$$\langle n \rangle \left( \frac{2\pi\hbar^2}{m k_B T_c} \right)^{3/2} = g_{3/2}(1)$$

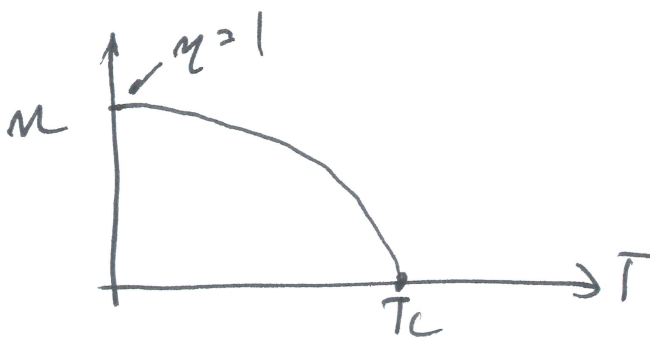
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below  $T_c$ :  $\langle n \rangle = n_0 + \frac{g_{3/2}(1)}{\lambda_T^3}$

$$\langle n \rangle = n_0 + \langle n \rangle \left( \frac{T}{T_c} \right)^{3/2}$$

$$\eta = \frac{n_0}{\langle n \rangle} = 1 - \left( \frac{T}{T_c} \right)^{3/2}$$

$\eta \rightarrow$  BEC fraction  $\rightarrow$  order parameter



at  $T=0 \Rightarrow \langle n \rangle = n_0$   
whole system is BEC

determine  $P$ - $V$  coexistence curve

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eliminate  $T$ :

$$P = \frac{k_B T}{\lambda^3} g_{5/2}(1) = \frac{k_B (m k_B)^{3/2} T^{5/2}}{(2\pi\hbar^2)^{3/2}} g_{5/2}(1)$$

$$\langle n \rangle = \frac{(m k_B)^{3/2} T^{3/2}}{(2\pi\hbar^2)^{3/2}} g_{3/2}(1)$$

$$T^{3/2} = \frac{(2\pi\hbar^2)^{3/2}}{(m k_B)^{3/2}} \frac{1}{g_{3/2}(1)} \langle n \rangle$$

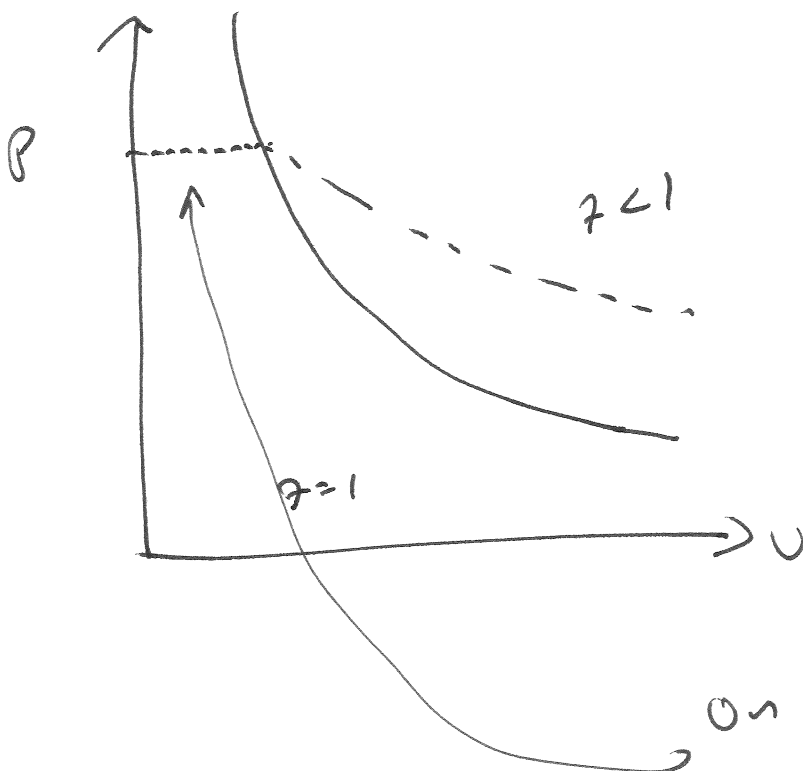
$$T = \frac{(2\pi\hbar^2)}{(m k_B)} \frac{\langle n \rangle^{2/3}}{[g_{3/2}(1)]^{2/3}}$$

$$P = \frac{k_B (m k_B)^{3/2}}{(2\pi\hbar^2)^{3/2}} \frac{(2\pi\hbar^2)^{5/2}}{(m k_B)^{5/2}} \frac{\langle n \rangle^{5/3}}{[g_{3/2}(1)]^{5/3}} g_{5/2}(1)$$

$\Downarrow$

$\Downarrow$

$$P_c \sim \frac{1}{V_c^{5/3}}$$



for isotherms

$$\frac{\partial P}{\partial V} \leq 0 \Rightarrow \text{stability}$$

in the BEC region

$$(T=1)$$

$$P = \frac{k_B T}{\lambda^3} g_{5/2}(1)$$

$\Downarrow$

$P$  does not depend  
on  $V \Rightarrow$  constant

calculate heat capacity / entropy

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$$dU = -SdT - PdV$$

$$\Rightarrow S = \left. \frac{\partial U}{\partial V} \right|_{T, \mu} = \left. \frac{\partial P}{\partial T} \right|_{V, \mu}$$

$$S = \frac{5}{2} \frac{k_B}{\lambda_T^{3/2}} g_{5/2}(z) + \frac{k_B T}{\lambda_T^3} g'_{5/2}(z) \frac{\partial z}{\partial T} \quad \boxed{z < 1}$$

$$\frac{\partial z}{\partial T} = \frac{\partial e^{\beta \mu}}{\partial T} = e^{\beta \mu} \left( -\frac{1}{k_B T^2} \right) = -\frac{z \beta \mu}{T}$$

$$S = \frac{5}{2} \frac{k_B}{\lambda_T^{3/2}} g_{5/2}(z) + \frac{k_B}{\lambda_T^3} g'_{5/2}(z) z \beta \mu$$

$$S = \frac{5}{2} \frac{k_B}{\lambda_T^{3/2}} g_{5/2}(z) - \frac{k_B}{\lambda_T^3} g_{3/2}(z) \ln z$$

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$$S = \begin{cases} \frac{5}{2} \frac{k_B}{\lambda_T^{3/2}} g_{5/2}(z) - k_B \ln z \ln z & z < 1 \\ \frac{5}{2} \frac{k_B}{\lambda_T^{3/2}} g_{5/2}(1) & z = 1 \end{cases}$$


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$T \left. \frac{\partial S}{\partial T} \right|_n \Rightarrow$  need derivative under condition  $n$ -fixed

derivation will need  $\frac{\partial z}{\partial T}$  with  $\ln z$  fixed

$$\ln z = \frac{1}{\lambda_T^3} g_{3/2}(z)$$

$$0 = +\frac{3}{2} \frac{1}{\lambda_T^3} T g_{3/2}(z) + \frac{1}{\lambda_T^3} z g'_{3/2}(z) \frac{1}{T} \frac{\partial z}{\partial T}$$

$$\frac{3}{2T} g_{3/2}(z) + g'_{3/2}(z) \frac{\partial \ln z}{\partial T} \Rightarrow \frac{\partial \ln z}{\partial T} = -\frac{3 g_{3/2}(z)}{2T g'_{3/2}(z)}$$

$$\left. \frac{\partial S}{\partial T} \right|_n = k_B \frac{15}{4} \frac{1}{\lambda_T^3} g_{5/2}(z) + k_B \frac{5}{2} \frac{1}{\lambda_T^3} g'_{3/2}(z) \frac{\partial z}{\partial T} \quad (12)$$

$$- k_B \langle n \rangle \frac{\partial \ln z}{\partial T}$$

$$= k_B \frac{15}{4} \frac{1}{\lambda_T^3} g_{5/2}(z) + k_B \frac{5}{2} \frac{1}{\lambda_T^3} g'_{3/2}(z) \frac{\partial \ln z}{\partial T}$$

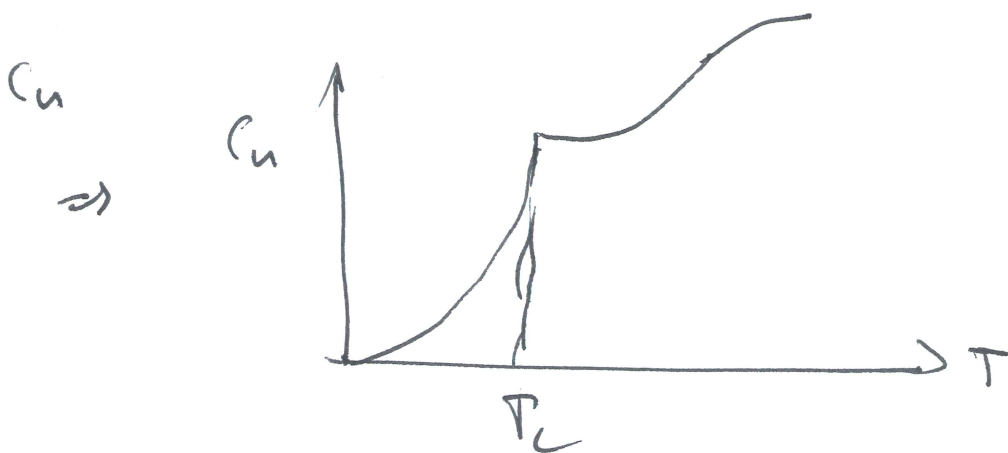
$$- k_B \langle n \rangle \frac{\partial \ln z}{\partial T}$$

$$= k_B \left( \frac{15}{4} \right) \frac{1}{\lambda_T^3} \frac{1}{T} g_{5/2}(z) + k_B \frac{5}{2} \langle n \rangle \frac{\partial \ln z}{\partial T}$$

$$= k_B \left( \frac{15}{4} \right) \frac{1}{\lambda_T^3} \frac{1}{T} g_{5/2}(z) + k_B \frac{9}{4T} \langle n \rangle \frac{g_{3/2}(z)}{g_{5/2}(z)}$$

$$C_n = T \left. \frac{\partial S}{\partial T} \right|_n = \begin{cases} k_B \left( \frac{15}{4} \right) \frac{1}{\lambda_T^3} g_{5/2}(z) - k_B \frac{9}{4} \langle n \rangle \frac{g_{3/2}(z)}{g_{5/2}(z)} & z < 1 \\ k_B \left( \frac{15}{4} \right) \frac{1}{\lambda_T^3} g_{5/2}(1) & z = 1 \end{cases}$$

$S \rightarrow 0$  as  $T \rightarrow 0 \Rightarrow$  obeys third law of thermodynamics



# Ideal Fermi-Dirac Gas

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- spin  $s = 1/2$ , mass  $m$ , box  $L$

$$Z_{FD}(\bar{T}, V, \mu) = \prod_i \left( \sum_{n_i=0}^1 e^{-\beta(\epsilon_i - \mu)} \right) \left( \sum_{n_i=0}^1 e^{-\beta(\epsilon_i - \mu/n_i)} \right)$$

$$= \prod_i (1 + e^{-\beta(\epsilon_i - \mu)})^2$$

- if spin  $\downarrow$   $3/2 \Rightarrow$  4 states

$$Z_{FD}(\bar{T}, V, \mu) = \prod_i (1 + e^{-\beta(\epsilon_i - \mu)})^4$$

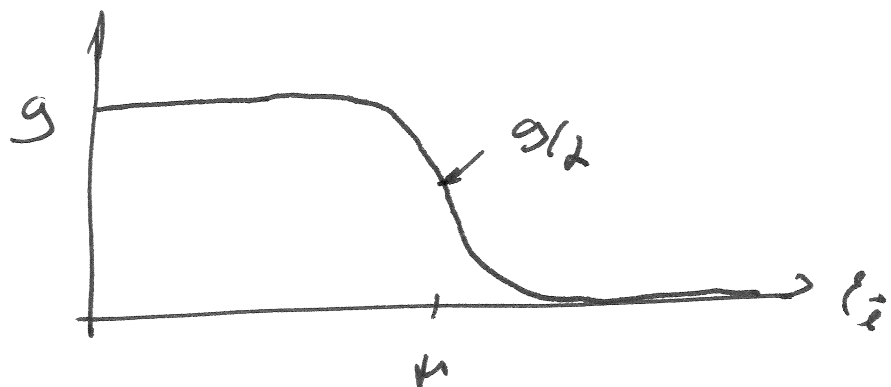
$$\Omega_{FD}(\bar{T}, V, \mu) = -k_B T \ln Z_{FD}(\bar{T}, V, \mu) = -k_B T g \sum_i \ln(1 + e^{-\beta(\epsilon_i - \mu)})$$

$$\langle N \rangle = - \frac{\partial \Omega_{FD}(\bar{T}, V, \mu)}{\partial \mu} = -k_B T g \sum_i \frac{e^{-\beta(\epsilon_i - \mu)}}{1 + e^{-\beta(\epsilon_i - \mu)}} (-\beta)$$

$$= g \sum_i \frac{1}{e^{\beta(\epsilon_i - \mu)} + 1}$$

$$\langle n_i \rangle = \frac{g_i}{e^{\beta(\epsilon_i - \mu)} + 1}$$

$\Rightarrow$  distribution:



change sum to integration

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$$-P_{FD}(\bar{T}, V, \mu) = -PV = -\frac{4\pi^2 g h^3 T^3 V}{(2\pi g)^3} \int_0^\infty p^2 dp \ln[1 + e^{-\beta(p^2/\mu)}]$$

$$\langle N \rangle = \frac{4\pi^2 g V}{(2\pi g)^3} \int_0^\infty p^2 dp \left( \frac{2}{e^{\beta(p^2/\mu)} + 1} \right)$$

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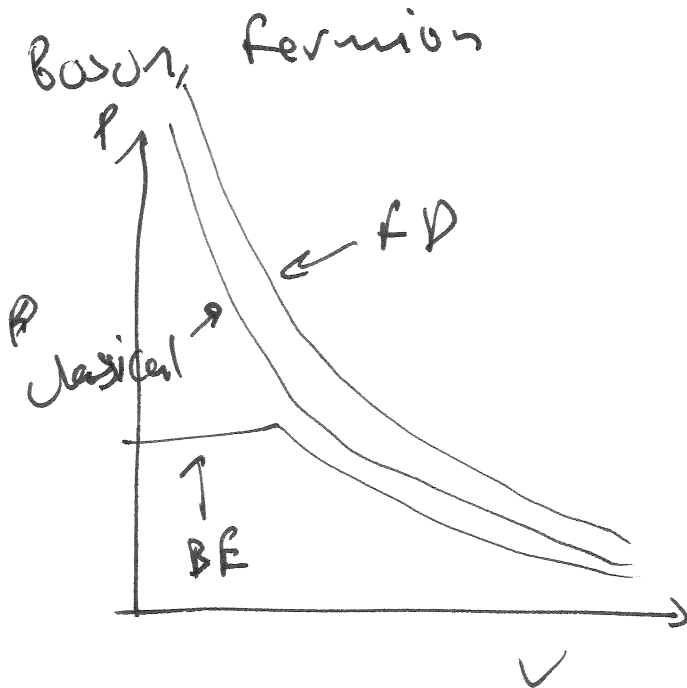
$$P = -\frac{\partial F_{FD}}{\partial V} = \frac{g h^3 T^3}{\lambda_T^3} f_{5/2}(\bar{z})$$

$$\langle N \rangle = \frac{\langle N \rangle}{V} = \frac{g}{\lambda_T^3} f_{3/2}(\bar{z})$$

$$f_{5/2}(\bar{z}) = \frac{4}{\sqrt{\pi}} \int_0^\infty x^2 dx \ln[1 + \bar{z} e^{-x^2}] = \sum_{\alpha=1}^{\infty} (-1)^{\alpha+1} \frac{\bar{z}^\alpha}{\alpha^{3/2}}$$

$$f_{3/2}(\bar{z}) = \bar{z} \frac{d}{d\bar{z}} f_{5/2}(\bar{z}) = \frac{4}{\sqrt{\pi}} \int_0^\infty x^2 dx \left( \frac{\bar{z}}{e^{\bar{z}^{-1} x^2} + 1} \right) = \sum_{\alpha=1}^{\infty} (-1)^{\alpha+1} \frac{\bar{z}^\alpha}{\alpha^{3/2}}$$

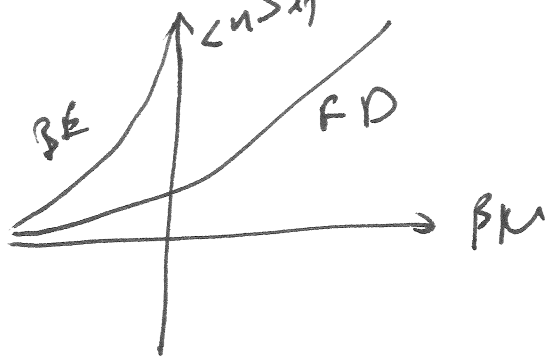
pressure vs. volume for ideal gases: classical,



BE  $\Rightarrow$  condensate does not contribute to pressure

$\langle n \rangle \lambda_T^3$

v.s.  $\mu$



BE  $\mu \rightarrow 0$  ( $T \rightarrow 0$ )

$\Rightarrow \langle n \rangle \lambda_T^3 \rightarrow \infty$

FD  $\mu \rightarrow \text{finite}$

$\omega T \rightarrow 0$

invert series

$\langle n \rangle \lambda_T^3 = \frac{5}{2} f_{3/2}(z)$

$f_{3/2}(z) = \sum_{n=1}^{\infty} (-1)^{n+1} \frac{z^n}{n^{3/2}}$

$z = \frac{\langle n \rangle \lambda_T^3}{g} + \frac{1}{2^{3/2}} \left( \frac{\langle n \rangle \lambda_T^3}{g} \right)^2 + \left( \frac{1}{2^2} - \frac{1}{3^{3/2}} \right) \left( \frac{\langle n \rangle \lambda_T^3}{g} \right)^3 + \dots$

as  $T \rightarrow \infty$   $z \rightarrow 0$

( $\beta \mu \Rightarrow$  large and negative)

as  $\beta \rightarrow 0$   $\mu \rightarrow -\infty$

$z \rightarrow \infty$  low  $T \Rightarrow z \rightarrow \infty$

$\beta \mu \rightarrow \infty$   $\beta \rightarrow \infty$

$\mu \rightarrow$  can remain finite

$f_{3/2}(z) = \frac{4}{\sqrt{\pi}} \int_0^{\infty} dx \frac{x^2}{(e^{x^2 - \nu} + 1)}$

$\gamma = x^2$

$x = \sqrt{\gamma}$

$dx = \frac{d\gamma}{2\sqrt{\gamma}}$

$= \frac{2}{\sqrt{\pi}} \int_0^{\infty} \frac{d\gamma}{2\sqrt{\gamma}} \frac{\sqrt{\gamma}}{(e^{\gamma - \nu} + 1)}$

$= \frac{2}{\sqrt{\pi}} \int_0^{\infty} \frac{d\gamma \sqrt{\gamma}}{(e^{\gamma - \nu} + 1)} = \frac{4}{3\sqrt{\pi}} \int_0^{\infty} d\gamma \gamma^{3/2} \frac{e^{\gamma - \nu}}{(e^{\gamma - \nu} + 1)^2}$

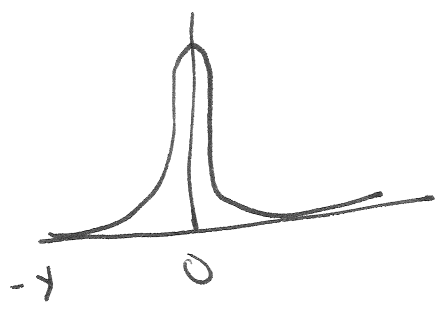
(integration by parts)

$t = \gamma - \nu$

$$\frac{\gamma}{3\sqrt{\pi}} \int_{-\nu}^{\infty} dt t^{3/2} \frac{e^{-t}}{(e^t + 1)^2}$$

$$t^{3/2} = \nu^{3/2} + \frac{3}{2} \nu^{1/2} t + \frac{3}{8} \nu^{1/2} t^2 + \dots$$

$$\approx \frac{\gamma}{3\sqrt{\pi}} \int_{-\nu}^{\infty} dt \frac{e^{-t}}{(e^t + 1)^2} \left[ \nu^{3/2} + \frac{3}{2} \nu^{1/2} t + \frac{3}{8} \nu^{1/2} t^2 + \dots \right]$$



$\nu = \beta \mu$

$$\frac{e^{-t}}{(e^t + 1)^2} \rightarrow \frac{e^{-\beta \mu}}{e^{\beta \mu}}$$

can extend integral to  $-\infty$

$$= \frac{\gamma}{3\sqrt{\pi}} \int_{-\infty}^{\infty} dt \frac{e^{-t}}{(e^t + 1)^2} \left[ \nu^{3/2} + \frac{3}{2} \nu^{1/2} t + \frac{3}{8} \nu^{1/2} t^2 + \dots \right]$$

$$I_n = \int_{-\infty}^{\infty} dt \frac{e^{-t}}{(e^t + 1)^2} t^n$$

$I_n = 0$  for  $n$  odd

$$I_n = \frac{(n-1)!}{(\ln 2)^{n-1}} \zeta(n) \quad n \text{ even}$$

Wilman's table function



$$\langle n \rangle \frac{\rho_T^3}{g} = \frac{4}{3\sqrt{\pi}} \left[ (\beta\mu)^{3/2} + \frac{\pi^2}{8} (\beta\mu)^{-1/2} + \dots \right] \quad (17)$$

$$T \rightarrow 0 \Rightarrow \mu(T=0) = \frac{\hbar^2}{2m} \left( \frac{6\sqrt{2} \langle n \rangle}{g} \right)^{2/3}$$

↓

Fermi energy

at  $T=0\text{K}$  Fermi energy is the maximum energy a particle can have

at low  $T$  particles will be excited to on the order of  $k_B T$

⇒ physics takes place near Fermi surface

$$\mu = \epsilon_F \left( 1 - \frac{\pi^2}{12} \left( \frac{k_B T}{\epsilon_F} \right)^2 + \dots \right)$$

can also calculate internal energy

$$U = \frac{3}{5} \langle n \rangle \epsilon_F \left( 1 + \frac{5\pi^2}{12} \left( \frac{k_B T}{\epsilon_F} \right)^2 + \dots \right)$$

$$C_V = \frac{\langle n \rangle \pi^2}{2} \frac{k_B T}{\epsilon_F}$$

specific heat is linear at low- $T$

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$\Rightarrow$  usually in metals electronic degrees of freedom contribute linearly