## In This Lecture:

$>$ Semiconductor Linear Optics: General Absorption, Emission Expressions

## 'The Electromagnetic Field [SI Units in action]

As you know, the machinery of QM works with potentials:
$\vec{E}(\stackrel{\rightharpoonup}{r}, t)=-\frac{\partial \vec{A}(\vec{r}, t)}{\partial t}-\vec{\nabla} \phi(\vec{r}, t)$,
$B(\vec{r}, t)=\vec{\nabla} \times \vec{A}(\vec{r}, t)$.
We shall work in the Coulomb gauge:
$\vec{\nabla} \cdot \vec{A}(\vec{r}, t)=0$,
$\phi(\vec{r}, t)=\frac{1}{4 \pi \varepsilon} \int d^{3} r^{\prime} \frac{\rho\left(\vec{r}^{\prime}, t\right)}{\left|\vec{r}-\vec{r}^{\prime}\right|}$.

However, the effect of the optical (EM) field is to cause certain transitions (interband, intraband) creating at most e-h pairs, therefore the local charge neutrality is always preserved,
i.e., $\quad \rho(\vec{r}, t)=0 \Rightarrow \phi(\vec{r}, t)=0$.

So, $\quad \vec{E}(\vec{r}, t)=-\frac{\partial \vec{A}(\vec{r}, t)}{\partial t}$.

Assume a monochromatic planewave EM field:
$\vec{A}(\vec{r}, t)=\hat{e} A_{0} \cos \left(\vec{k}_{o p} \cdot \vec{r}-\omega t\right) \Rightarrow \vec{E}(\vec{r}, t)=-\hat{e} \omega A_{0} \sin \left(\vec{k}_{o p} \cdot \vec{r}-\omega t\right)$.

The associated optical intensity is given by:
$I=|\langle\vec{S}(\vec{r}, t)\rangle|=\left|\left\langle\frac{\vec{E} \times \vec{B}}{\mu_{0}}\right\rangle\right|=\frac{n_{r} c \varepsilon_{0} \omega^{2} A_{0}^{2}}{2}$.

In the presence of EM field the canonical momentum operator, $\vec{p}=-i \hbar \vec{\nabla}$ of an electron $(e<0)$ is replaced by the kinetic momentum, $\vec{p} \rightarrow \vec{p}-e \vec{A}$, so that,

$$
\begin{aligned}
H & =\frac{1}{2 m_{0}}(\vec{p}-e \vec{A})^{2}+e \phi+V(\vec{r}), \\
& =\frac{p^{2}}{2 m_{0}}-\frac{e}{2 m_{0}}(\vec{p} \cdot \vec{A}+\vec{A} \cdot \vec{p})+\frac{e^{2}}{2 m_{0}} A^{2}+e \not \phi+V(\vec{r}),
\end{aligned}
$$

Note that, $\quad \frac{e}{2 m_{0}} \vec{p} \cdot \vec{A}=e \frac{\vec{A} \cdot \vec{p}}{2 m_{0}}-i \frac{e \hbar}{2 m_{0}} \vec{\nabla} / \vec{A}$,
Coulomb Gauge

$$
H=\frac{p^{2}}{2 m_{0}}-\frac{e}{m_{0}} \vec{A} \cdot \vec{p}+\frac{e^{2}}{2 m_{0}} A^{2}+V(\vec{r}), \quad \begin{aligned}
& \text { Compare their typical } \\
& \text { relative magnitudes }
\end{aligned}
$$

For most practical optical field intensities: $|e \vec{A}| \ll|\vec{p}| \approx \frac{\hbar \pi}{a_{0}}$,
so that $\frac{e A}{p} \sim 10^{-5}$ for a $1 \mathrm{~W} / \mathrm{cm}^{2}$ photon density of a 1 eV energy illumination.

Thus neglecting the $A^{2}$ term leaves us with

$$
H=\underbrace{\frac{p^{2}}{2 m_{0}}+V(\vec{r})}_{H_{0}} \underbrace{-\frac{e}{m_{0}} \vec{A} \cdot \vec{p}, \quad \begin{array}{c}
\text { Radiation perturbation } \\
\text { (time-dependent) }
\end{array}}_{H^{\prime}(t)}
$$

The nicety of the radiation perturbation is that the matrix elements of the perturbation already takes part in the $\vec{k} \cdot \vec{p}$ band structure techniques.

Terminology: The result we have obtained is also termed as the so-called velocity gauge. An alternative expression follows in the next page. However, in the rest of our analysis we shall carry on from the expression on this page.

## Electric Dipole Approximation (Length Gauge)

Note that the Hamiltonian as expressed in the previous page depends on the specific gauge (Coulomb) we have chosen. Another form for the radiation perturbation that is in common use is given by:

$$
H^{\prime}(t)=-e \stackrel{\rightharpoonup}{r} \cdot \stackrel{\rightharpoonup}{E}(\stackrel{\rightharpoonup}{r}, t) \quad \ldots \ldots \ldots \ldots . . . . .
$$

It can be shown that in the long wavelength of the EM wave ( $k_{\text {op }}$ small) the electric dipole appoximation agrees with the A•p results. As a matter of fact, even though $\mathbf{r}$ operator becomes problematic for infinite solids, it can be handled and moreover, the r-E expression has its own virtues from computational perspectives of the linear and nonlinear susceptibilies; see, for instance Aversa-Sipe PRB 5214636 (1995).

## Further Reflections

On Seeing the Light


There's some that swear by whisky, There's some that swear by rye, There's some that swear by A.p And others by $e \phi$.

But those that swear by whisky Will now and then take rye, But those that swear by A.p Will never use $e \phi$.

If those that swear by whisky
Would know the reason why
All those who swear by A.p
Would never use $e \phi$.
Then ask the $e \phi$ people.
Their answer won't be guarded:
"All those who swear by A.p
Are definitely retarded."
In the A.p man's heaven
There's a place forever bright
For repentant $e \phi$ sinners
Who at last have seen the light.
B. K. Ridley (1992)

## Absorption, Spontaneous/Stimulated Emission

Due to time-dependent perturbation, we need to make use of Fermi's golden rule. The resultant transition rate among the unperturbed initial/final states is given by

$$
\left.W(i)=\frac{2 \pi}{\hbar} \sum_{f}\left|\langle f| H^{\prime}\right| i\right\rangle\left.\right|^{2} \delta\left(E_{f}-E_{i} \mp \hbar \omega\right) \quad \boldsymbol{H}^{\prime}=-\frac{\boldsymbol{e}}{\boldsymbol{m}_{0}} \overrightarrow{\boldsymbol{A}} \cdot \overrightarrow{\boldsymbol{p}}
$$

We need to switch (temporarily) from our semiclassical approach to the photon picture; (i.e., second quantized EM field). The aim is to capture the stimulated/spontaneous emission processes.

Interestingly, the relation between absorption and stim/spon emission was first proposed by Einstein w/o even invoking standard QM, leading to Einstein's $A$ and $B$ coefficients; see Chuang for this alternative route.

$$
\text { Absorption: }|f\rangle=\left|k_{f}, n_{p h}-1\right\rangle, \quad|i\rangle=\left|k_{i}, n_{p h}\right\rangle \gg \text { Emission: }|f\rangle=\left|k_{f}, n_{p h}+1\right\rangle
$$

## Second Quantization



Wave nature is manifested only if wavelength is comparable to or smaller than distances over which potential energy changes significantly.


Particle nature is manifested when either/or:

- wave intensity is very low so that the number of quanta approaches $\sim 1$
- Interactions with matter involve exchange of single quantum.

Figure 6.13: A conceptual picture of first and second quantization. The second quantization allows us to describe classical waves as quantum mechanical particles.


Figure 9.3: (a) A schematic of an absorption process where a photon is absorbed (destroyed) and the energy and momentum of the electron is altered; (b) the emission of a photon where a photon is created.

When second-quantized, the vector potential's amplitude becomes:

$$
A_{0}=\sqrt{\frac{\hbar}{2 \omega \varepsilon V}}\left(b_{\star}^{\dagger}+b\right)
$$

Photon creation/annihilation operators; monochromatic (mode index suppressed)

Properties: $\quad b^{\dagger}\left|n_{p h}\right\rangle=\sqrt{n_{p h}+1}\left|n_{p h}+1\right\rangle$, $b\left|n_{p h}\right\rangle=\sqrt{n_{p h}}\left|n_{p h}-1\right\rangle$

Absorption:

$$
\begin{aligned}
\langle i| \vec{A} \cdot \vec{\nabla}|f\rangle & =\sqrt{\frac{\hbar}{2 \omega \varepsilon}}\left\langle k_{f}, n_{p h}-1\right|\left(b^{\dagger}+b\right) \hat{e} \cdot \vec{\nabla}\left|k_{i}, n_{p h}\right\rangle \\
& =\sqrt{\frac{\hbar n_{p h}}{2 \omega \varepsilon}}\left\langle k_{f}\right| \hat{e} \cdot \vec{\nabla}\left|k_{i}\right\rangle
\end{aligned}
$$

Emission:

$$
\begin{aligned}
\langle i| \vec{A} \cdot \vec{\nabla}|f\rangle & =\sqrt{\frac{\hbar}{2 \omega \varepsilon}}\left\langle k_{f}, n_{p h}+1\right|\left(b^{\dagger}+b\right) \hat{e} \cdot \vec{\nabla}\left|k_{i}, n_{p h}\right\rangle \\
& =\sqrt{\frac{\hbar\left(n_{p h}+1\right)}{2 \omega \varepsilon}}\left\langle k_{f}\right| \hat{e} \cdot \vec{\nabla}\left|k_{i}\right\rangle
\end{aligned}
$$

From now on we shall treat absorption and emission separately because of the different settings of the processes.

For the absorption, assume an incoming photon with momentum $\hbar k_{o p}$ and energy $\hbar \omega$ to be absorbed by the se/c via all possible initial/final states as long as energy/momentum conservations are satisfied.


Absorption rate
per unit volume of se/c

Polarization dependence of absorption is in this matrix element

## Absorption Rate $\rightarrow$ Absorption Coefficient

Start with the continuity equation for photons:

In the steady state $d / d t \rightarrow 0$ and the form that defines the absorption coefficient $\alpha$ is given by

$$
n_{p h}(x)=n_{0} e^{-\alpha x}
$$

Inserting this form into the continuity equation yields:

$$
W_{a b s}=\alpha v n_{p h} \quad \text { or } \quad \alpha=\frac{W_{a b s}}{v n_{p h}}
$$

## Absorption Coefficient of Some se/c

| It is the absorption |
| :---: |
| coefficient that is |
| provided in the literature |



Ref: Singh

## Kramers-Kronig Relations (KKR)

$>$ KKR is based on the principle of causality: the response to an applied field such as the polarization (effect) cannot precede the applied field (cause) $>$ Mathematically, it follows from the integral representation of the step function (Heaviside fn) over the complex plane

$$
\underset{\text { Dielectric Function }}{\text { Complex }} \varepsilon(\omega)=\varepsilon^{\prime}(\omega)+i \varepsilon^{\prime \prime}(\omega)
$$



$$
\left.\begin{array}{l}
\varepsilon^{\prime}(\omega)-\varepsilon_{0}=\frac{2}{\pi} \boldsymbol{P}\left(\int_{0}^{\infty} \frac{\omega^{\prime} \varepsilon^{\prime \prime}(\omega)}{\omega^{\prime 2}-\omega^{2}} d \omega^{\prime}\right), \\
\varepsilon^{\prime \prime}(\omega)=-\frac{2 \omega}{\pi} \boldsymbol{P}\left(\int_{0}^{\infty} \frac{\varepsilon^{\prime}\left(\omega^{\prime}\right)-\varepsilon_{0}}{\omega^{\prime 2}-\omega^{2}} d \omega^{\prime}\right)
\end{array}\right\} \begin{gathered}
\text { written as Hilbert } \\
\text { transform pairs }
\end{gathered}
$$

For a non-magnetic ( $\mu_{0}$ ) homogeneous isotropic medium (such as bulk cubic xtal) the wave vector of an EM wave is given by:

$$
\begin{aligned}
& k(\omega)=\beta+i \alpha \\
& k(\omega)=\omega \sqrt{\mu_{0} \varepsilon(\omega)} \\
& k(\omega)=\frac{\omega}{c} \sqrt{\varepsilon_{r}(\omega)} ; \text { where } \varepsilon_{r}=\varepsilon / \varepsilon_{0} \\
& \beta^{2}-\alpha^{2}=\left(\frac{\omega}{c}\right)^{2} \varepsilon_{r}^{\prime} \\
& 2 \alpha \beta=\left(\frac{\omega}{c}\right)^{2} \varepsilon_{r}^{\prime \prime} \square \begin{array}{l}
\text { Absorption coef. is proportional to the imaginary part } \\
\text { of the diel. fn. which in turn effects on the index of } \\
\text { refraction through the KKR }
\end{array}
\end{aligned}
$$

## Absorption vs Emission Rates

In the absorption expression, if we are considering the rate at which a photon is absorbed, this involves summing over any possible electronic levels, as long as the energy conservation is obeyed. One may instead, be interested in the problem where a single electron with (xtal) momentum $k_{i}$ is scattered into a final state with momentum $k_{f}$. In this case the final states sum would be over all photon states that can cause such transitions and would involve photon density of states. These two approaches are illustrated in the following figure.


Approaches are linked by $\rho(\hbar \omega)$ and $N(E)$

## Continue with the Emission Rate

The rate of an electron $\hbar k_{i}$ to emit any photon and reach a state with momentum $\hbar k_{f}$ is given by

$$
W_{e m}=\frac{2 \pi e^{2}}{\hbar m_{0}^{2}}\left(\frac{\hbar\left(n_{p h}+1\right)}{2 \omega \varepsilon}\right) \sum_{\substack{\text { photon } \\ \text { states }}}\left|\int \psi_{f}^{*} \hat{e} \cdot \vec{p} e^{-i \vec{k}_{o p} \cdot \vec{r}} \psi_{i} d^{3} r\right|^{2} \delta\left(E_{i}-E_{f}-\hbar \omega\right)
$$

Notice that the emission term can be rewritten as stimulated and spontaneous emission terms as

$$
\begin{gathered}
W_{e m}=W_{s t}+W_{\text {spon }} \\
W_{s t}=\frac{2 \pi e^{2}}{\hbar m_{0}^{2}} \frac{\hbar n_{p h}}{2 \omega \varepsilon_{\substack{\text { photon } \\
\text { states }}}\left|\int \psi_{f}^{*} \hat{e} \cdot \vec{p} e^{-i \vec{k}_{o p} \cdot \vec{r}} \psi_{i} d^{3} r\right|^{2} \delta\left(E_{i}-E_{f}-\hbar \omega\right)} \\
W_{\text {spon }}=\frac{2 \pi e^{2}}{\hbar m_{0}^{2}} \frac{\hbar}{2 \omega \varepsilon_{\substack{\text { photon } \\
\text { states }}}\left|\int \psi_{f}^{*} e^{-i \vec{k}_{o p} \cdot \cdot} \psi_{i} d^{3} r\right|^{2} \delta\left(E_{i}-E_{f}-\hbar \omega\right)}
\end{gathered}
$$

## Spontaneous vs Stimulated Emission

The stimulated emission is due to the initial photons present in the system and the emitted photons maintain phase coherence with the initial photons. The spontaneous emission comes from the perturbations due to the vacuum state (i.e., $n_{p h}=0$ ) energy fluctuations and the emitted photons are incoherent with no phase relationship. The difference between these two kinds of processes is the key to understanding the differences between a light emitting diode and the laser diode

## Spontaneous em.



SPONTANEOUS EMTSSION: Responsible for light emission in LEDs

- Electron-hole recombination in the absence of photons
- Outcoming photons arg ingaherent. ie, have randomphases
$:$ Electron-hole recombination lifetime is limited by ~l 1 is. :


## Stimulated em.



Stmulated kmission: Responsible for light emission in laser diodes

- Electron-hole recombination in the presence of other photons
- Photons produced are coherent, i.e. have the same phase
- Electron-hole recombination lifetime is:

$$
\frac{1}{\tau_{\text {stuve }}}=\frac{n_{\text {ph }}}{\tau_{\text {spum }}}
$$

: $n_{p h}:$ photon number

