## In This Lecture:

$>$ Interband Transitions in Bulk Se/c

* Momentum Matrix Element
* Polarization dependence
$>$ Interband Transitions in Quantum Wells
> Intraband Transitions in Bulk \& QWs


## Interband Transitions in Bulk Se/c

As the photon momentum is negligible compared to electronic crystal momenta, the transitions are (almost)
vertical ( $k_{\text {op }}=0$ )
[See, next page]


Figure 9.5: The positions of the electron and hole energies at vertical $k$-values. The electron and hole energies are determined by the photon energy and the carrier masses. Since the photon momentum is negligible the transitions are vertical.

## Momentum

Predominantly we are interested in transitions between CB and VB so that $i \rightarrow v, f \rightarrow c$

## Matrix Elements

$$
\hat{e} \cdot \vec{p}=\hat{e} \cdot \int \psi_{c, \bar{k}_{c}}^{*} e^{i \bar{k}_{o p} \cdot \bar{r}} \vec{p} \psi_{v, \bar{k}_{v}} d^{3} r
$$

where

$$
\psi_{c, \bar{k}_{c}}(\vec{r})=\frac{1}{\sqrt{V}} e^{i \vec{k}_{c} \cdot \vec{r}} u_{c, \bar{k}_{c}}(\vec{r}) ; \quad \text { similarly for } \psi_{v, \bar{k}_{v}}(\vec{r})
$$

vanishes when $k_{\text {op }}=0$ due to
Bloch fn orthogonality

$$
\begin{aligned}
\vec{p}_{c v} & =\int_{c, \overrightarrow{k_{c}}}^{*}(\vec{r}) e^{-i \vec{k}_{c} \cdot \vec{r}} e^{i \vec{k}_{o p} \cdot \vec{r}} \hbar k_{v} e^{i \vec{k}_{v} \cdot \vec{r}} u_{v, \vec{k}_{v}}(\vec{r}) \frac{d^{3} r}{V} \\
& +\int u_{c, \vec{k}_{c}}^{*}(\vec{r}) e^{-i \vec{k}_{c} \cdot \vec{r}} e^{i \vec{k}_{o p} \cdot \vec{r}}\left(\frac{\hbar}{i} \vec{\nabla} u_{v, \bar{k}_{v}}(\vec{r})\right) e^{i \vec{k}_{v} \cdot \vec{r}} \frac{d^{3} r}{V}
\end{aligned}
$$

Treat slowly-varying (envelopes) and the cell-periodic parts separately

$$
\vec{p}_{c v}=\delta_{\vec{k}_{c}, \vec{k}_{v}+\vec{k}_{o p}} \int_{\Omega} u_{c, \bar{k}_{c}}^{*}(\vec{r})\left(\frac{\hbar}{i} \stackrel{\nabla}{\nabla} u_{v, \bar{k}_{v}}(\vec{r})\right) \frac{d^{3} r}{\Omega}
$$

## Electric dipole forbidden transitions

When certain $\boldsymbol{p}_{c v}$ transition matrix element vanishes (due to some symmetry reason etc.) this is termed as a electric dipole-forbidden-transition.
In this case higher-order contributions such as electric quadrupole and magnetic dipole transitions become important. Compared to the electric dipole transitions they are reduced in strength by a factor of (lattice constant/wavelength of light) ${ }^{2}$, that requires very high frequencies (UV to X-rays)...

## Polarization Dependence

Momentum-matrix parameter:

$$
P_{x}=\langle i S| p_{x}|X\rangle=\langle i S| p_{y}|Y\rangle=\langle i S| p_{z}|Z\rangle=\frac{m_{0}}{\hbar} P
$$

CB to HH Transitions:

$$
\begin{aligned}
& \langle i S \uparrow| \vec{p}\left|\frac{3}{2}, \frac{3}{2}\right\rangle=-\frac{P_{x}}{\sqrt{2}}(\hat{x}+i \hat{y}), \\
& \langle i S \downarrow| \vec{p}\left|\frac{3}{2}, \frac{3}{2}\right\rangle=0 \\
& \langle i S \downarrow| \vec{p}\left|\frac{3}{2},-\frac{3}{2}\right\rangle=\frac{P_{x}}{\sqrt{2}}(\hat{x}-i \hat{y}), \\
& \langle i S \uparrow| \vec{p}\left|\frac{3}{2},-\frac{3}{2}\right\rangle=0,
\end{aligned}
$$

## WATCH OUT:

No coupling of the z-polarized light between CB \& HH

## CB to LH Transitions:

$$
\begin{aligned}
& \langle i S \uparrow| \vec{p}\left|\frac{3}{2}, \frac{1}{2}\right\rangle=P_{x} \sqrt{\frac{2}{3}} \hat{z} \\
& \langle i S \downarrow| \vec{p}\left|\frac{3}{2}, \frac{1}{2}\right\rangle=-\frac{P_{x}}{\sqrt{6}}(\hat{x}+i \hat{y}), \\
& \langle i S \downarrow| \vec{p}\left|\frac{3}{2},-\frac{1}{2}\right\rangle=P_{x} \sqrt{\frac{2}{3}} \hat{z} \\
& \langle i S \uparrow| \vec{p}\left|\frac{3}{2},-\frac{1}{2}\right\rangle=\frac{P_{x}}{\sqrt{6}}(\hat{x}-i \hat{y})
\end{aligned}
$$

## Reflections on the polarization dependence

$>$ For a cubic xtal what differentiates $z$ from $x$ or $y$ ?
$>$ Recall that in defining the expansion basis vectors we assumed electron wavevector to be along $z$ direction
$>$ For that reason we are also using the z-projection of the spin
$>$ Does that give enough support for singling out $z$ from $x$ or $y$ direction?
$>$ After all that's just for the sake of formulation, say a convention
$\rightarrow$ Away from $\mathbf{k}=0 \mathrm{HH} \& \mathrm{LH}$ become mixed
$>$ So only at $\mathbf{k}=0$ we could talk about such a selectivity
$>$ But at $\mathbf{k}=0$ we lose any sense of direction of the $k$-vector!

## To Remind you the LK Hamiltonian

$$
\overline{\overline{\mathbf{H}}}^{\mathrm{LK}}=-\left[\begin{array}{cccccc}
P+Q & -S & R & 0 & -S / \sqrt{2} & \sqrt{2} R \\
-S^{+} & P-Q & 0 & R & -\sqrt{2} Q & \sqrt{3 / 2} S \\
R^{+} & 0 & P-Q & S & \sqrt{3 / 2} S^{+} & \sqrt{2} Q \\
0 & R^{+} & S^{+} & P+Q & -\sqrt{2} R^{+} & -S^{+} / \sqrt{2} \\
-S^{+} / \sqrt{2} & -\sqrt{2} Q^{+} & \sqrt{3 / 2} S & -\sqrt{2} R & P+\Delta & 0 \\
\sqrt{2} R^{\oplus} & \sqrt{3 / 2} S^{+} & \sqrt{2} Q^{+} & -S / \sqrt{2} & 0 & P+\Delta
\end{array}\right]
$$

complex conjugate

$$
\text { where }\left\{\begin{array}{l}
P=\frac{\hbar^{2} \gamma_{1}}{2 m_{0}}\left(k_{x}^{2}+k_{y}^{2}+k_{z}^{2}\right) \\
Q=\frac{\hbar^{2} \gamma_{2}}{2 m_{0}}\left(k_{x}^{2}+k_{y}^{2}-2 k_{z}^{2}\right) \\
R=\frac{\hbar^{2}}{2 m_{0}}\left[-\sqrt{3} \gamma_{2}\left(k_{x}^{2}-k_{y}^{2}\right)+i 2 \sqrt{3} \gamma_{3} k_{x} k_{y}\right] \\
S=\frac{\hbar^{2} \gamma_{3}}{m_{0}} \sqrt{3}\left(k_{x}-\mathrm{i} k_{y}\right) k_{z}
\end{array}\right.
$$

## Averaging over the polarization for bulk

$>$ These considerations suggest us to consider unpolarized light
>Equivalently we shall consider electron wavevector to point along a general direction and average the matrix element over the solid angle

Let the electron wavevector to be along a direction $(\theta, \Phi)$ :

$$
\mathbf{k}=k \sin \theta \cos \phi \hat{x}+k \sin \theta \sin \phi \hat{y}+k \cos \theta \hat{z}
$$

For illustration consider CB-HH transition:

$$
\left.\left.\left|\hat{e} \cdot \mathbf{p}_{c v}\right|^{2} \equiv\langle | \hat{e} \cdot \mathbf{M}_{c-h h}\right|^{2}\right\rangle=\underbrace{\frac{1}{4 \pi} \int\left|\hat{x} \cdot \mathbf{M}_{c-h h}\right|^{2} \sin \theta \mathrm{~d} \theta \mathrm{~d} \phi}_{\begin{array}{c}
\text { averaging over } \\
\text { the solid angle }
\end{array}}
$$

CB: $\quad\left|\mathrm{i} S \downarrow^{\prime}\right\rangle$ and $\left|\mathrm{i} S \uparrow^{\prime}\right\rangle$
$\left(\left|\frac{3}{2}, \frac{3}{2}\right\rangle^{\prime}=\frac{-1}{\sqrt{2}}\left|\left(X^{\prime}+i Y^{\prime}\right) \uparrow^{\prime}\right\rangle\right.$

$$
\left.=\frac{-1}{\sqrt{2}} \right\rvert\,(\cos \theta \cos \phi-\mathrm{i} \sin \phi) X
$$

$$
+(\cos \theta \sin \phi+\mathrm{i} \cos \phi) Y-\sin \theta Z\rangle\left|\uparrow^{\prime}\right\rangle
$$

$$
\left|\frac{3}{2},-\frac{3}{2}\right\rangle^{\prime}=\frac{1}{\sqrt{2}}\left|\left(X^{\prime}-\mathrm{i} Y^{\prime}\right) \downarrow^{\prime}\right\rangle
$$

$$
\left.=\frac{1}{\sqrt{2}} \right\rvert\,(\cos \theta \cos \phi+\mathrm{i} \sin \phi) X
$$

$$
+(\cos \theta \sin \phi-i \cos \phi) Y-\sin \theta Z\rangle\left|\downarrow^{\prime}\right\rangle
$$

Note that for ease of calculation we keep the spin parts in the new (rotated) coordinate system...

$$
\begin{aligned}
& \begin{aligned}
\left\langle\mathrm{i} S \uparrow^{\prime}\right| \mathbf{p}\left|\frac{3}{2}, \frac{3}{2}\right\rangle^{\prime}=- & {[(\cos \theta \cos \phi-\mathrm{i} \sin \phi) \hat{x}} \\
& +(\cos \theta \sin \phi+\mathrm{i} \cos \phi) \hat{y}-\sin \theta \hat{z}] \frac{P_{x}}{\sqrt{2}}
\end{aligned} \\
& \begin{aligned}
\left\langle\mathrm{i} S \downarrow^{\prime}\right| \mathbf{p}\left|\frac{3}{2},-\frac{3}{2}\right\rangle^{\prime}= & {[(\cos \theta \cos \phi+\mathrm{i} \sin \phi) \hat{x}} \\
& +(\cos \theta \sin \phi-\mathrm{i} \cos \phi) \hat{y}-\sin \theta \hat{z}] \frac{P_{x}}{\sqrt{2}}
\end{aligned} \\
& \begin{array}{l}
\left\langle\mathrm{i} S \uparrow^{\prime}\right| \mathbf{p}\left|\frac{3}{2},-\frac{3}{2}\right\rangle^{\prime}= \\
\left\langle\mathrm{i} S \downarrow^{\prime}\right| \mathbf{p}\left|\frac{3}{2}, \frac{3}{2}\right\rangle^{\prime}=
\end{array}
\end{aligned}
$$

Consider, for instance optical transition from the CB of one spin, say $<\mathrm{i} S \uparrow^{\prime} \mid$ to either of the HH bands $\left|\frac{3}{2}, \frac{3}{2}\right\rangle^{\prime}\left|\frac{3}{2},-\frac{3}{2}\right\rangle^{\prime}$; one of them is already zero

## Bulk Momentum Matrix Element for Unpolarized Light

$$
\begin{aligned}
\left.\left.\left|\hat{e} \cdot \mathbf{p}_{c v}\right|^{2} \equiv\langle | \hat{e} \cdot \mathbf{M}_{c-h h}\right|^{2}\right\rangle & =\frac{1}{4 \pi} \int\left|\hat{x} \cdot \mathbf{M}_{c-h h}\right|^{2} \sin \theta \mathrm{~d} \theta \mathrm{~d} \phi \\
& =\frac{1}{4 \pi} \int_{0}^{\pi} \sin \theta \mathrm{d} \theta \int_{0}^{2 \pi} \mathrm{~d} \phi\left(\cos ^{2} \theta \cos ^{2} \phi+\sin ^{2} \phi\right) \frac{P_{x}^{2}}{2} \\
& =\frac{1}{3} P_{x}^{2} \equiv M_{b}^{2}
\end{aligned}
$$

where $\quad M_{b}^{2}=\frac{1}{3} P_{x}^{2}=\frac{m_{0}^{2}}{3 \hbar^{2}} P^{2}$
Kane's parameter, (not a surprise)

$$
=\left(\frac{m_{0}}{m_{e}^{*}}-1\right) \frac{m_{0} E_{g}\left(E_{g}+\Delta\right)}{6\left(E_{g}+\frac{2}{3} \Delta\right)}
$$

Alternatively, an energy parameter $E_{p}$ can be defined as:

$$
E_{p}=\frac{2 m_{0}}{\hbar^{2}} P^{2}, \text { so that } M_{b}=\frac{m_{0}}{6} E_{p}
$$

## The other polarizations, spin, and LH band

Same result $M_{b}^{2}$ is obtained for

* For $\hat{e}=\hat{y}$ or $\hat{e}=\hat{z} \quad$ (cubic symmetry)
* For the other spin component of the CB, $\left\langle i S \downarrow \downarrow^{\prime}\right|$
* For the transition between the LH band (per spin),

$$
\left.\left.\left|\left\langle i S \downarrow^{\prime}\right| e x\right| \frac{3}{2}, \frac{1}{2}\right\rangle\left.^{\prime}\right|^{2}+\left|\left\langle i S \downarrow^{\prime}\right| e x\right| \frac{3}{2},-\frac{1}{2}\right\rangle\left.^{\prime}\right|^{2}
$$

## Joint Density of States (also called reduced DOS)

This is an important piece that appears inside total transition rate expressions

## Single Parabolic Band DOS: <br> $$
N_{m}(E)=\sum_{\bar{k} \in \mathbb{I N}^{\mathrm{Bz}}} \sum_{\sigma} \delta\left(E-E_{m}(\vec{k})\right)
$$

For a parabolic band: $E-E_{c}=\frac{\hbar^{2} k^{2}}{2 m_{\text {dos }}^{*}}$

$$
N_{m}(E)=\sqrt{2} \frac{\left(m_{d o s}^{*}\right)^{3 / 2} \sqrt{E-E_{c}}}{\pi^{2} \hbar^{3}},
$$

## Joint DOS of CB-VB:



Between two parabolic CB and VB: $\hbar \omega-E_{g}=\frac{\hbar^{2} k^{2}}{2} \underbrace{\left(\frac{1}{m_{e}^{*}}+\frac{1}{m_{h}^{*}}\right)}_{\frac{1}{m_{r}^{*}}}$

$$
N_{c v}(\hbar \omega)=\sum_{\vec{k} \in \mathrm{I}^{\mathrm{L}} \mathrm{BZ}} \sum_{\sigma} \delta\left(E_{v}(\vec{k})-E_{c}(\vec{k})+\hbar \omega\right)
$$

$$
N_{c v}(\hbar \omega)=\sqrt{2} \frac{\left(m_{r}^{*}\right)^{3 / 2} \sqrt{\hbar \omega-E_{g}}}{\pi^{2} \hbar^{3}}
$$

## Absorption Rate (Final Expression)

With all these ingredients the bulk absorption rate for unpolarized light becomes:

$$
\begin{aligned}
& W_{a b s}=\frac{\pi e^{2} \hbar n_{p h}}{m_{0}^{2} \hbar \omega \varepsilon}\left(2 M_{b}^{2}\right) N_{c v}(\hbar \omega) \\
& \\
& N_{c v}(\hbar \omega)=\sqrt{2} \frac{\left(m_{r}^{*}\right)^{3 / 2} \sqrt{\hbar \omega-E_{g}}}{\pi^{2} \hbar^{3}}
\end{aligned}
$$

## Radiative e-h Recombination Time: Emission

In the case of interband recombination rate of an e with a hole at the same $\mathbf{k}$ state, we integrate over all possible photon states


$$
W_{\mathrm{em}}=\frac{\pi e^{2} \hbar}{m_{o}^{2} \hbar \omega \boldsymbol{\epsilon}}\left(n_{p h}+1\right)\left|\boldsymbol{a} \cdot \boldsymbol{p}_{\mathrm{if}}\right|^{2} \rho_{a}(\hbar \omega)
$$

$$
\rho(\hbar \omega)=\frac{2 \omega^{2}}{2 \pi^{2} \hbar v^{3}} \quad 3 \mathrm{D} \text { total photon DOS }
$$

For $n_{p h}=0, \quad W_{e m} \rightarrow W_{s p o n}$

Associated e-h radiative recombination time is $\tau_{0}=\frac{1}{W_{\text {spon }}}$

## Interband Transitions in Quantum Wells

transitions between subbands derived from different bulk bands


## Subband Wavefunctions



3D to 2D: Optical transitions are affected in two ways
> Form of JDOS
> Momentum matrix element; anisotropy is now genuine

## Momentum Matrix Element in QWs

In going from 3D to 2D:

$$
\begin{aligned}
& p_{\text {if }}^{\text {3D }}=\frac{1}{V} \int e^{i\left(\mathbf{k}_{k}-\mathbf{k}_{n}\right) \cdot \mathbf{r}\left\langle u_{v}^{v}\right| p_{a}\left|u_{c}\right\rangle d^{3} r} \\
& \rightarrow p_{\mathrm{ff}}^{2 \mathrm{D}}=\frac{1}{A W} \sum_{\nu} \underbrace{\left(g_{v}^{\nu / m}\left|g_{c}^{n}\right\rangle\right.} \underbrace{e^{i\left(\mathbf{k}_{e}-\mathbf{k}_{h}\right) \cdot \rho}\left\langle u_{v}^{\nu m}\right| p_{c}\left|u_{c}\right\rangle d^{2} \rho} \\
& \text { env. fn. overlap in-plane overlap } \\
& \text { along growth dir. }
\end{aligned}
$$

[ Other term, $p_{a}$ acting on $g_{c}^{n}(z)$ leaves $\left\langle u_{v}^{v m} \mid u_{c}\right\rangle=0$ at the same $\vec{k}$ state]
Unlike 3D, polarization dependence exists in 2D
Notation $\left\{\begin{array}{l}\text { TE (to growth axis): Electric field in QW plane } \\ \text { TM (to growth axis): Electric field along growth axis }\end{array}\right.$


## Let the QW growth axis be $\boldsymbol{z}$ axis

TE (Optical electric field in $x y$ plane)
Optical dipole matrix element is averaged over the azimuthal angle From both $\mathbf{H H}$ bands to $\langle\mathrm{i} S \uparrow$ †

$$
\left.\left.\left|\hat{e} \cdot \mathbf{p}_{c u}\right|^{2} \equiv\langle | \hat{e} \cdot \mathbf{M}_{c-h n}\right|^{2}\right\rangle=\frac{1}{2 \pi} \int_{0}^{2 \pi} \mathrm{~d} \hat{\phi}\left|\hat{x} \cdot \mathbf{M}_{c-h h}\right|^{2}
$$

## Same results for

 the other CB spins not considered$$
\begin{aligned}
& =\frac{1}{2 \pi} \int_{0}^{2 \pi} \mathrm{~d} \phi\left(\cos ^{2} \theta \cos ^{2} \phi+\sin ^{2} \phi\right) \frac{P_{x}^{2}}{2} \\
& =\frac{3}{4}\left(1+\cos ^{2} \theta\right) M_{b}^{2}
\end{aligned}
$$

From both LH bands to $\langle\mathrm{i} S \downarrow 1$

$$
\left.\left.\left.\left.\langle | \hat{e} \cdot \mathbf{M}_{c-l h}\right|^{2}\right\rangle=\left.\frac{1}{2 \pi} \int_{0}^{2 \pi} \mathrm{~d} \phi\left(\left|\left\langle\mathrm{i} S \downarrow^{\prime}\right| p_{x}\right| \frac{3}{2}, \frac{1}{2}\right\rangle\right|^{2}+\left|\left\langle\mathrm{i} S \downarrow^{\prime}\right| p_{x}\right| \frac{3}{2},-\frac{1}{2}\right\rangle\left.\right|^{2}\right)
$$

Same results for the other CB spins not considered

$$
\begin{aligned}
& =\left(\frac{2}{3} \sin ^{2} \theta\left\langle\cos ^{2} \phi\right\rangle+\frac{1}{6} \cos ^{2} \theta\left\langle\cos ^{2} \phi\right\rangle+\frac{1}{6}\left\langle\sin ^{2} \phi\right\rangle\right) P_{x}^{2} \\
& =\left[\sin ^{2} \theta+\frac{1}{4}\left(\cos ^{2} \theta+1\right)\right] M_{b}^{2} \\
& =\left(\frac{5}{4}-\frac{3}{4} \cos ^{2} \theta\right) M_{b}^{2}
\end{aligned}
$$

## TM (Optical electric field along $z$ axis)

$$
\begin{aligned}
\left.\left.\langle | \hat{e} \cdot \mathbf{M}_{c-h h}\right|^{2}\right\rangle & =\frac{1}{2 \pi} \int_{0}^{2 \pi} \mathrm{~d} \phi\left|\hat{z} \cdot \mathbf{M}_{c-h h}\right|^{2}=\frac{3}{2} \sin ^{2} \theta M_{b}^{2} \\
\left.\left.\langle | \hat{e} \cdot \mathbf{M}_{c-h h}\right|^{2}\right\rangle & \left.\left.=\left.\frac{1}{2 \pi} \int_{0}^{2 \pi} \mathrm{~d} \phi\left(|\langle\mathrm{i} S \downarrow| e z| \frac{3}{2}, \frac{1}{2}\right\rangle\right|^{2}+|\langle\mathrm{i} S \downarrow| e z| \frac{3}{2},-\frac{1}{2}\right\rangle\left.\right|^{2}\right) \\
& =\left(\frac{1}{6} \sin ^{2} \theta+\frac{2}{3} \cos ^{2} \theta\right) P_{x}^{2} \\
& =\frac{1+3 \cos ^{2} \theta}{2} M_{b}^{2}
\end{aligned}
$$

Table 9.1 Summary of the Momentum Matrix Elements in Parabolic
$\underline{\text { Band Model }\left(\left|\hat{e} \cdot \mathbf{p}_{c e}\right|^{2}=|\hat{e} \cdot \mathbf{M}|^{2}\right)}$
Bulk $\left|\hat{x} \cdot \mathbf{p}_{c e}\right|^{2}=\left|\hat{y} \cdot \mathbf{p}_{c c}\right|^{2}=\left|\hat{z} \cdot \mathbf{p}_{c v}\right|^{2}=M_{b}^{2}=\frac{m_{0}}{6} E_{p}$

## Quantum Well

TE Polarization $(\hat{e}=\hat{x}$ or $\hat{y}) \quad$ TM Polarization $(\hat{e}=\hat{z})$
$\left.\left.\left.\langle | \hat{e} \cdot \mathbf{M}_{c-h h}\right|^{2}\right\rangle=\left.\frac{3}{4}\left(1+\cos ^{2} \theta\right) M_{b}^{2} \quad\langle | \hat{e} \cdot \mathbf{M}_{c-h h}\right|^{2}\right\rangle=\frac{3}{2} \sin ^{2} \theta M_{b}^{2}$
$\left.\left.\left.\langle | \hat{e} \cdot \mathbf{M}_{c-l h}\right|^{2}\right\rangle=\left.\left(\frac{5}{4}-\frac{3}{4} \cos ^{2} \theta\right) M_{b}^{2} \quad\langle | \hat{e} \cdot \mathbf{M}_{c-l h}\right|^{2}\right\rangle=\frac{1}{2}\left(1+3 \cos ^{2} \theta\right) M_{b}^{2}$
Conservation Rule
Sum $\left.\left.\left.\left.\langle | \hat{x} \cdot \mathbf{M}_{c-h}\right|^{2}\right\rangle+\left.\langle | \hat{y} \cdot \mathbf{M}_{c-h}\right|^{2}\right\rangle+\left.\langle | \hat{z} \cdot \mathbf{M}_{c-h}\right|^{2}\right\rangle=3 M_{b}^{2},(h=h h$ or $l h)$
Rules $\left.\left.\left.\langle | \hat{e} \cdot \mathbf{M}_{c-h}\right|^{2}\right\rangle+\left.\langle | \hat{e} \cdot \mathbf{M}_{c-l h}\right|^{2}\right\rangle=2 M_{b}^{2}$

## Back to Absorption Rate in QWs

JDOS in 2D: $\begin{aligned} & \frac{N_{c v}^{2 D}(\hbar \omega)}{W}=\frac{m_{r}^{*}}{\pi \hbar^{2} W} \sum_{n m}\left\langle g_{v}^{m} \mid g_{c}^{n}\right\rangle \theta\left(\boldsymbol{E}_{\mathrm{nm}}-\hbar \omega\right) \\ & \boldsymbol{E}_{\mathrm{nm}}=\boldsymbol{E}_{\mathrm{g} \pi \mathrm{p}}+\boldsymbol{E}_{c}^{n}+\boldsymbol{E}_{v}^{m}\end{aligned}$

$$
\alpha(\hbar \omega)=\frac{\pi e^{2} \hbar}{m_{o}^{2} c n_{r} \epsilon_{o}} \frac{1}{(\hbar \omega)}\left|\boldsymbol{a} \cdot \boldsymbol{p}_{\mathrm{if}}\right|^{2} \frac{N_{2 D}(\hbar \omega)}{W} \sum_{n_{\mathrm{r}} \lambda_{i}} f_{n m} \theta\left(\boldsymbol{E}_{m n}-\hbar \omega\right)
$$

Observe that even-odd parity transitions are not allowed due to vanishing of this overlap


Figure 9.7: Calculated absorption coefficient in a $100 \AA \mathrm{GaAs}^{2} / \mathrm{Al}_{0.3} \mathrm{Ga}_{0.7}$ As quantum well structure for in-plane polarized light. The HH transition is about three times stronger than the LH transition in this polarization. In a real material excitonic transition dominate near the bandedges as disucssed in the next chapter.

## Indirect Interband Transitions in Bulk

Common Indirect Se/c: $\mathrm{Si}, \mathrm{Ge}, \mathrm{C}, \mathrm{AlAs}, \mathrm{GaP}, \mathrm{AIP}, \mathrm{SiC}, \mathrm{AIN}(\mathrm{zb})$


With photon energies smaller than the direct band gap intermediate transitions can occur since energy need not be conserved

$$
W_{i f}(\boldsymbol{k})=\frac{2 \pi}{\hbar} \int_{f}\left\{\left|M_{\mathrm{em}}\right|^{2}+\left|M_{\mathrm{abs}}\right|^{2}\right\} \delta\left(\boldsymbol{E}_{f}-\boldsymbol{E}_{i}\right) \frac{d^{3} k}{(2 \pi)^{3}}
$$

Pathways which require phonon emission/absorption

## Form of the matrix elements:

$$
M_{\mathrm{abs}}=\frac{\frac{\left.\left|\langle c, \boldsymbol{k}+\boldsymbol{q}| H_{\mathrm{cp}}^{\mathrm{abs}}\right| \boldsymbol{c}, \boldsymbol{k}\right\rangle \mid}{} \frac{\left(\left.|c, \boldsymbol{k}| H_{\mathrm{ph}}^{\mathrm{abs}}|v, \boldsymbol{k}|\right|^{2}\right.}{\left(\boldsymbol{E}_{g \mathrm{r}}-\hbar \omega\right)^{2}}}{M_{\mathrm{em}}}=\frac{\left.\left.\left|\langle c, \boldsymbol{k}-\boldsymbol{q}| H_{\mathrm{ep}}^{\mathrm{em}}\right| c, \boldsymbol{k}\right\rangle| |\langle c, \boldsymbol{k}| H_{\mathrm{ph}}^{\mathrm{em}} \mid v, \boldsymbol{k}\right)\left.\right|^{2}}{\left(\boldsymbol{E}_{g \Gamma}-\hbar \omega\right)^{2}} \text { direct optical transitions }
$$

e-phonon scattering matrix elements due to optical phonon intervalley scattering with the associated matrix element:
$D_{i j}$ : Deformation potential

$$
M_{q}^{2}=\frac{\hbar D_{i j}^{2}}{2 \rho V \omega_{i j}}\left\{\begin{array}{l}
n\left(\omega_{i j}\right) \\
n\left(\omega_{i j}\right)+1
\end{array}\right\} \rightarrow \text { abs. } \begin{aligned}
& \rho: \text { Mass density } \\
& \omega_{i j}: \text { Intervalley phonon frequency } \\
& n\left(\omega_{i j}\right): \text { phonon occupancy (BE distr.) }
\end{aligned}
$$

## For parabolic bands, the absorption rate results in:

\# equivalent


Figure 9.10: Absorption coefficient of Si and Ge. Also shown is absorption coefficient for amorphous silicon which is almost like a direct gap semiconductor, since $k$-selection is not applicable.

$$
\begin{aligned}
& \text { valleys } \\
& \begin{aligned}
W_{\mathrm{abs}( }(\hbar \omega) & =\frac{M_{\mathrm{ph}}^{2} D_{i j}^{2} J_{v}\left(m_{c} m_{v}\right)^{3 / 2}}{8 \pi^{2}\left(\boldsymbol{E}_{g \Gamma}-\hbar \omega\right)^{2} \hbar^{6} \rho \omega_{i j}} \\
& \times\left[n\left(\omega_{i j}\right)\left(\hbar \omega-\boldsymbol{E}_{g \mathbf{k}^{\prime}}+\hbar \omega_{\mathrm{ij}}\right)^{2}\right. \\
& \left.+\left\{n\left(\omega_{i j}\right)+1\right\}\left(\hbar \omega-\boldsymbol{E}_{g \mathbf{k}^{\prime}}-\hbar \omega_{i j}\right)^{2}\right]
\end{aligned} \\
& \text { Photon-related } \quad M_{\mathrm{ph}}^{2}=\frac{e^{2} \hbar n_{\mathrm{ph}}\left|\boldsymbol{\alpha} \cdot \boldsymbol{p}_{\mathrm{if}}\right|^{2}}{2 m_{0}^{2} \epsilon \omega}
\end{aligned}
$$

## Note the contrast in $W_{\text {abs }}$

Direct Bandgap: $\left(\boldsymbol{H} \boldsymbol{w}-\boldsymbol{E}_{g}\right)^{1 / 2}$
Indirect Bandgap: $\left(\hbar \omega-\boldsymbol{E}_{\mathrm{th}}\right)^{2}$
In amorphous se/c, $k$-conservation requirement is relaxed (no periodicity, xtal momentum not a good quantum label) This results in higher absorption coefficient

## Intraband Transitions in Bulk Se/c

$>$ As each band at a $k$-state is single-valued $1^{\text {st }}$ order vertical intraband transitions are not possible
$>$ Intraband transitions must involve some second
 mechanism (phonon, ionized imp, defects...) to ensure momentum conservation
> Intraband transitions are also known as free carrier absorption and are effective in the cladding
 layers of lasers

## Drude Model (to explain free carrier absorption)

$m^{*} \ddot{x}+m^{*} \gamma \dot{x}+m^{*} \omega_{0}^{2}=e E_{0} \cos (\omega t)$
w/o scattering no net energy xfer; $e$ 's oscillate back and forth within the band

By introducing a scattering mechanism, energy gained by the e in one cycle will be partially lost in the form of, say phonon emission by the electron.

$$
\begin{aligned}
\alpha(\hbar \omega) & \propto \frac{1}{\omega^{2}} \\
& \propto \frac{1}{\mu}
\end{aligned}
$$

If the mobility is large (weak scattering) absorption coefficient becomes small

## Intraband Transitions in Quantum Wells


$>$ Since a number of subbands may originate from the same bulk band, certain inter-subband transitions (CB1-CB2) may be termed as intraband transitions in QWs
$>$ Such inter-subband transitions have great importance for far infrared detectors and forms the basis of Quantum Cascade Lasers

$$
\begin{aligned}
& \psi^{1}(\boldsymbol{k}, z)= \\
& \psi^{2}(\boldsymbol{k}, z)=\frac{g^{1}(z)}{g^{2}(z)} e^{i \mathbf{k} \cdot \rho} e^{i \mathrm{k} \cdot \rho} u_{n \mathbf{k}}^{1}(r) \\
& u_{n \mathbf{k}}^{2}(\boldsymbol{r})
\end{aligned}
$$

## Momentum Matrix Element:

$$
\boldsymbol{p}_{\mathrm{if}}=-\frac{i \hbar}{W} \int g^{2 *}(z) e^{-i \mathbf{k} \cdot \rho} \boldsymbol{a} \cdot \nabla g^{1}(z) e^{i \mathbf{k} \cdot \rho} d^{2} \rho d z
$$

If the polarization lies on the QW plane, then due to the orthogonality of the remaining envelope parts

$$
\left(g^{1}, g^{2}\right), p_{\mathrm{if}}=0
$$

$>$ Thus for EM wave polarized in the plane of the QW, inter-subband transition rate is zero (This can be relaxed under strong mixing of the cell-periodic parts as in the VB.)
$>$ For EM wave polarized along the QW growth axis (say z), we get

$$
\boldsymbol{p}_{\mathrm{if}}=\frac{-i \hbar}{W} \int g^{2 *}(z) \hat{z} \frac{\partial}{\partial z} g^{1}(z) d z \quad \Longleftrightarrow \quad\left|\boldsymbol{p}_{\mathrm{if}}\right| \approx \frac{\hbar}{W}
$$

Brings $g^{1}$ to the same parity with $g^{2}$

